

# Can we hope to find gravitational waves from the early cosmos?

THEP day, May 8th, 2025

**Carlo Tasillo,**  
**Uppsala University**

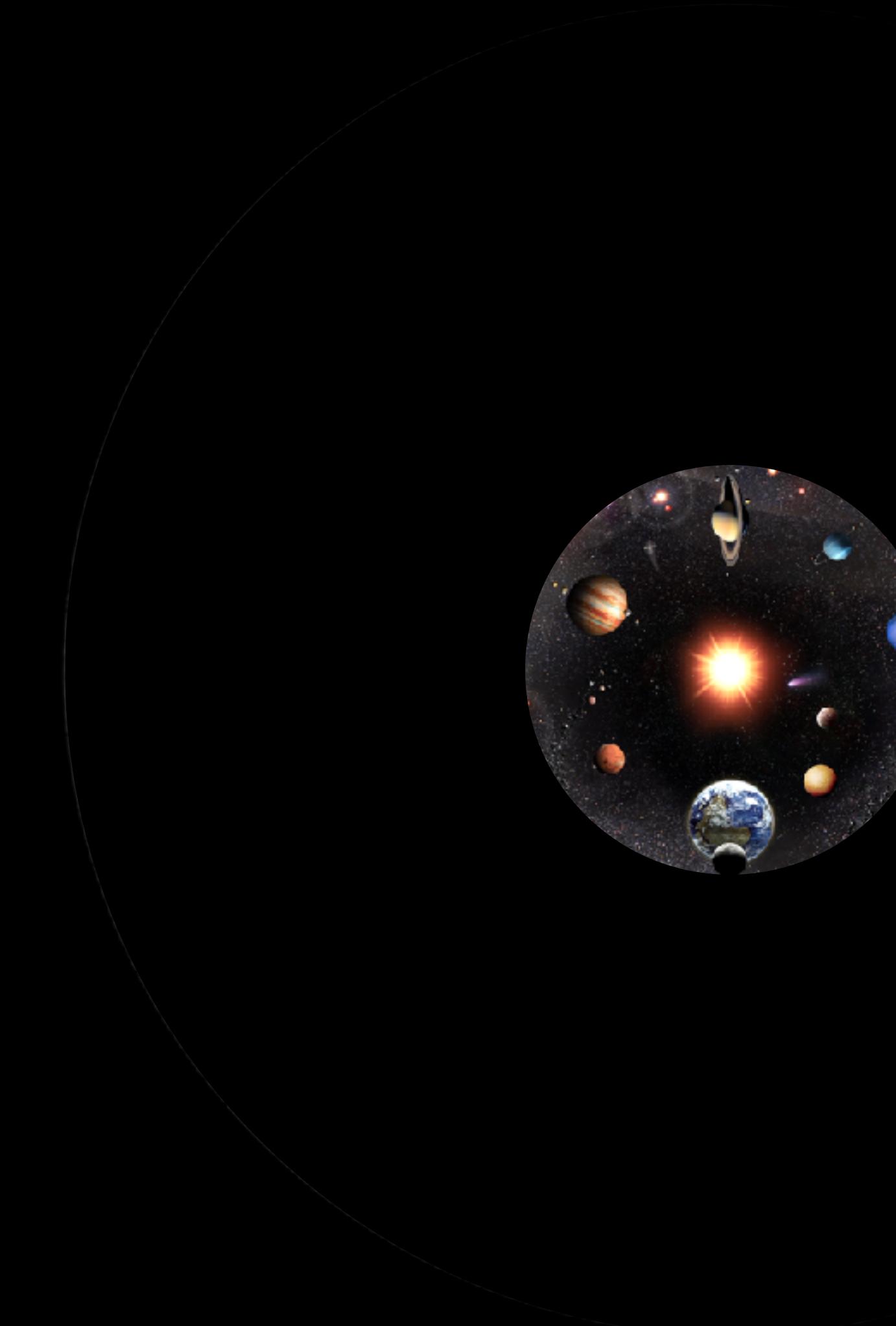
Based on ongoing work in collaboration with

- Torsten Bringmann, Felix Kahlhöfer, Thomas Konstandin, Jonas Matuszak and Kai Schmidt-Hoberg in Oslo, KIT, CERN-TH and DESY
- Safa, André, Rikard and Stefano in our THEP group



UPPSALA  
UNIVERSITET

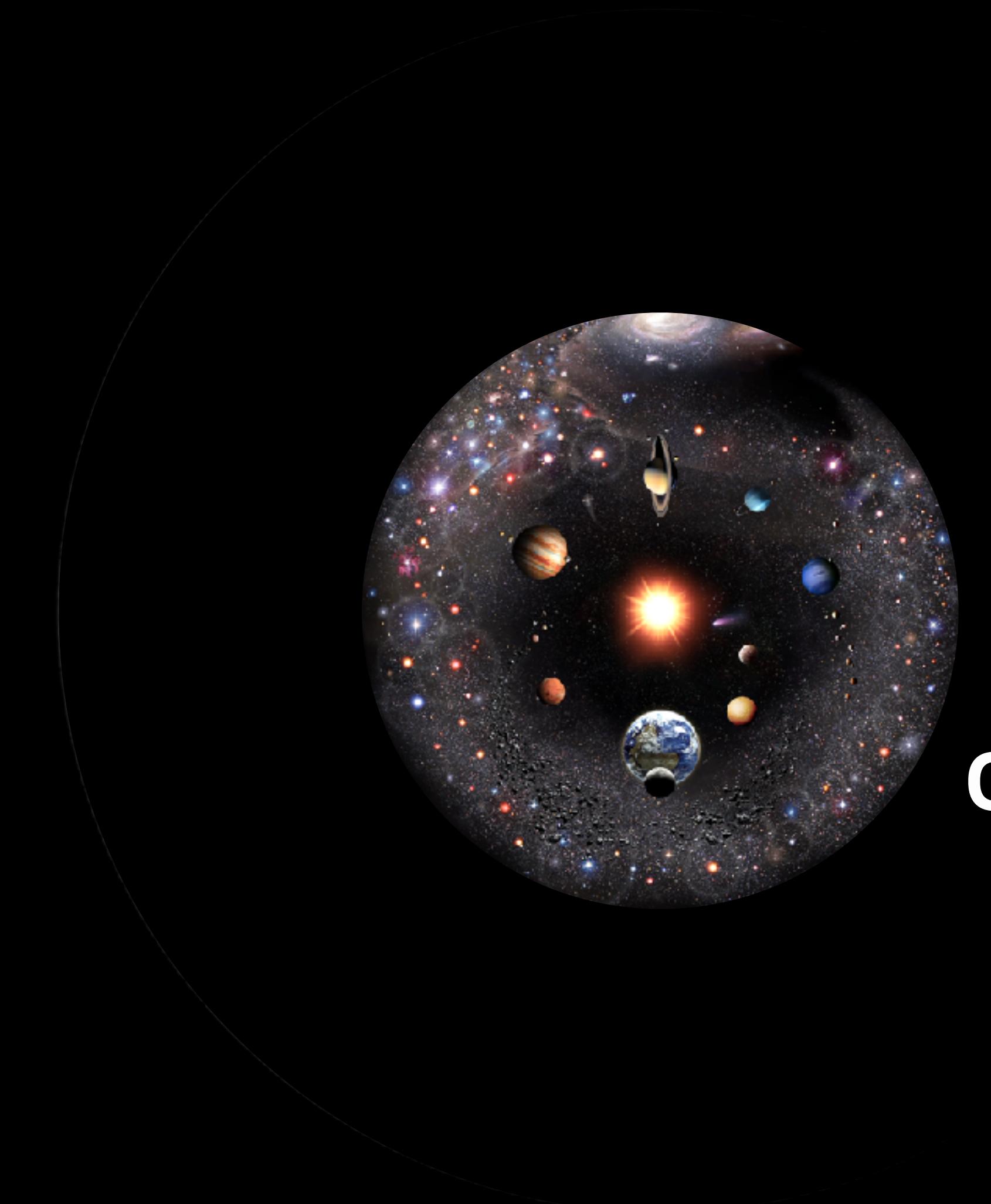
# The observable universe



**Our Solar System**

PABLO  
CARLOS  
BUDASSI

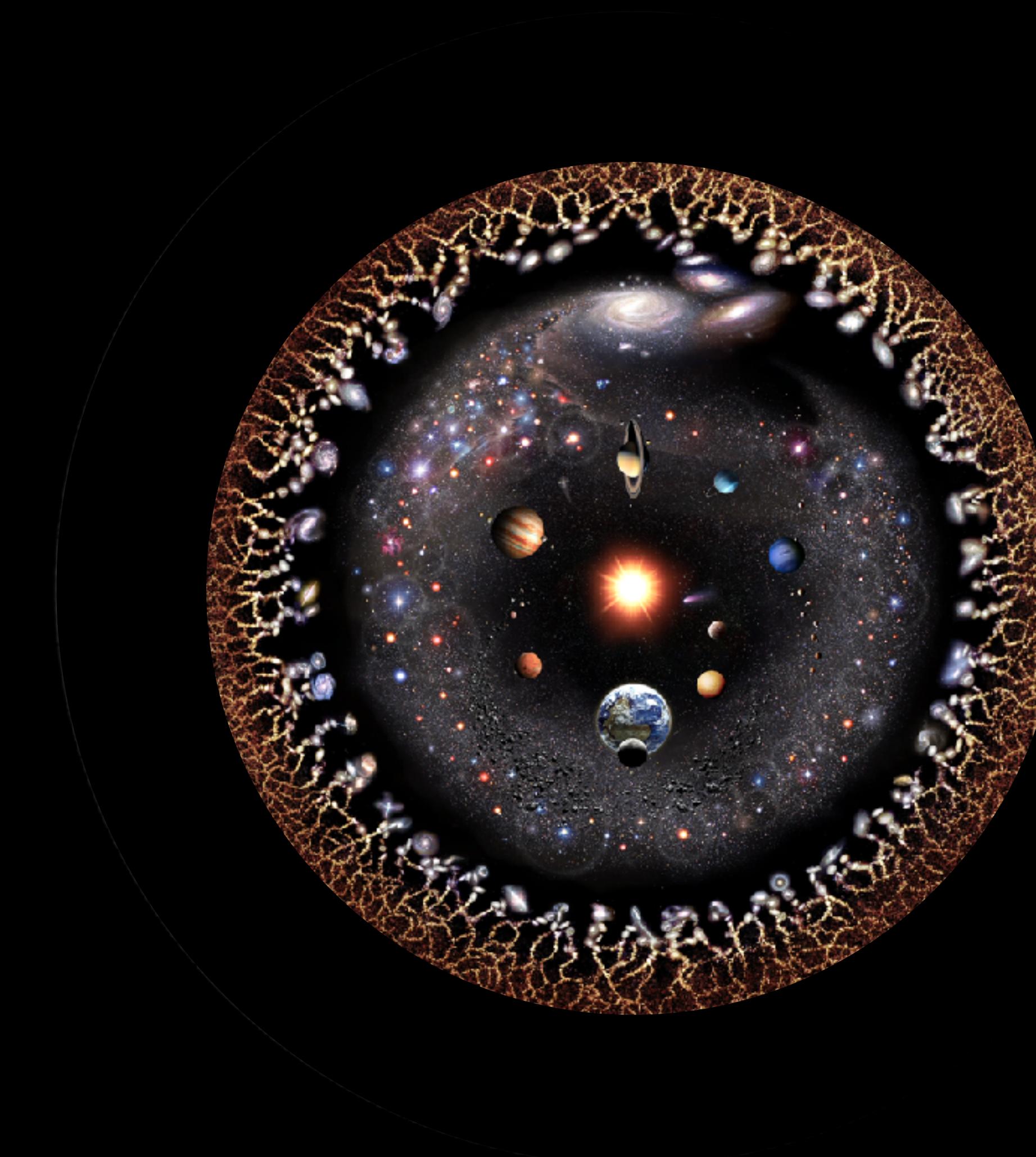
# The observable universe



Our galaxy

PABLO  
CARLOS  
BUDASSI

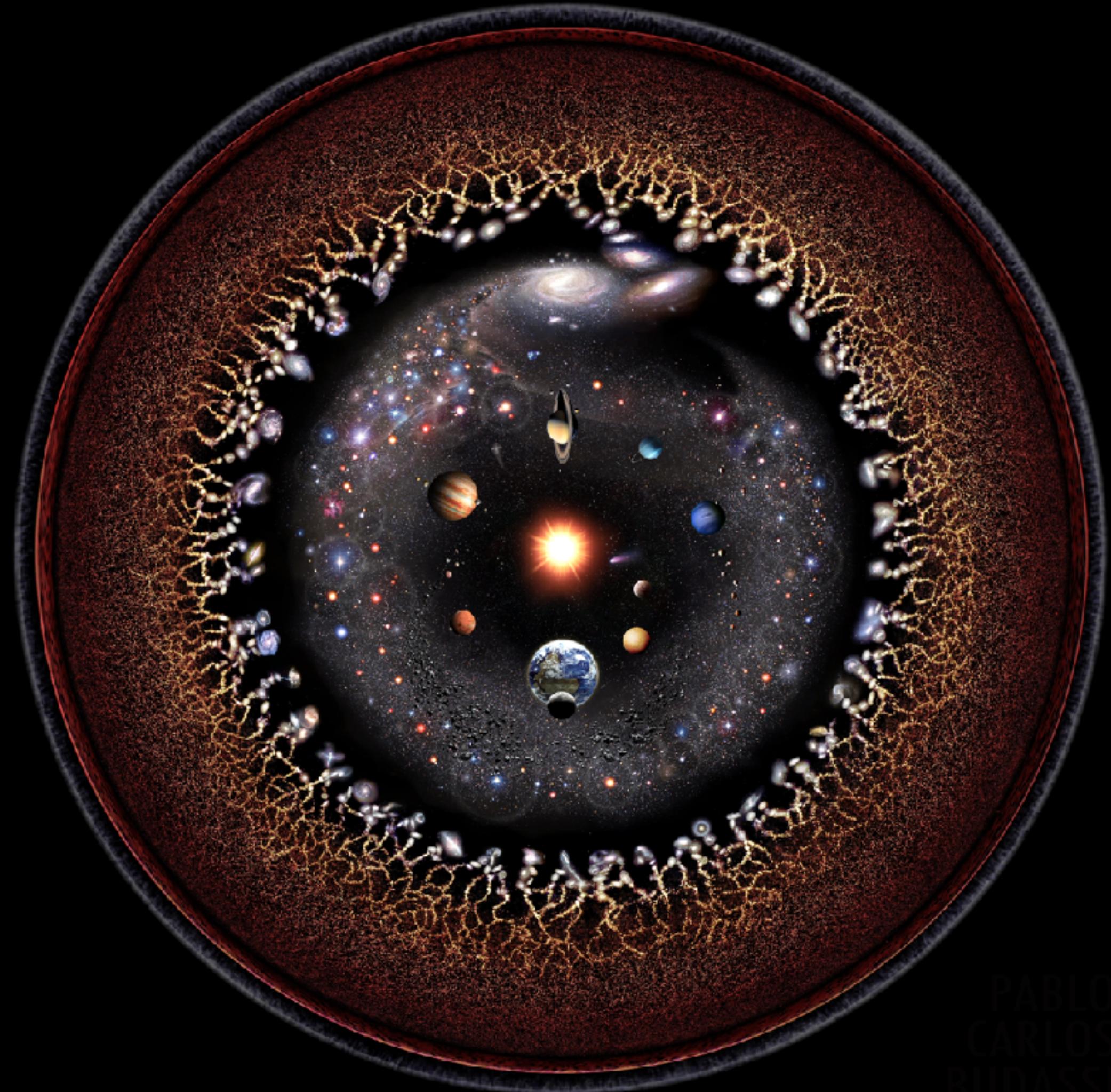
# The observable universe



Other galaxies

PABLO  
CARLOS  
BUDASSI

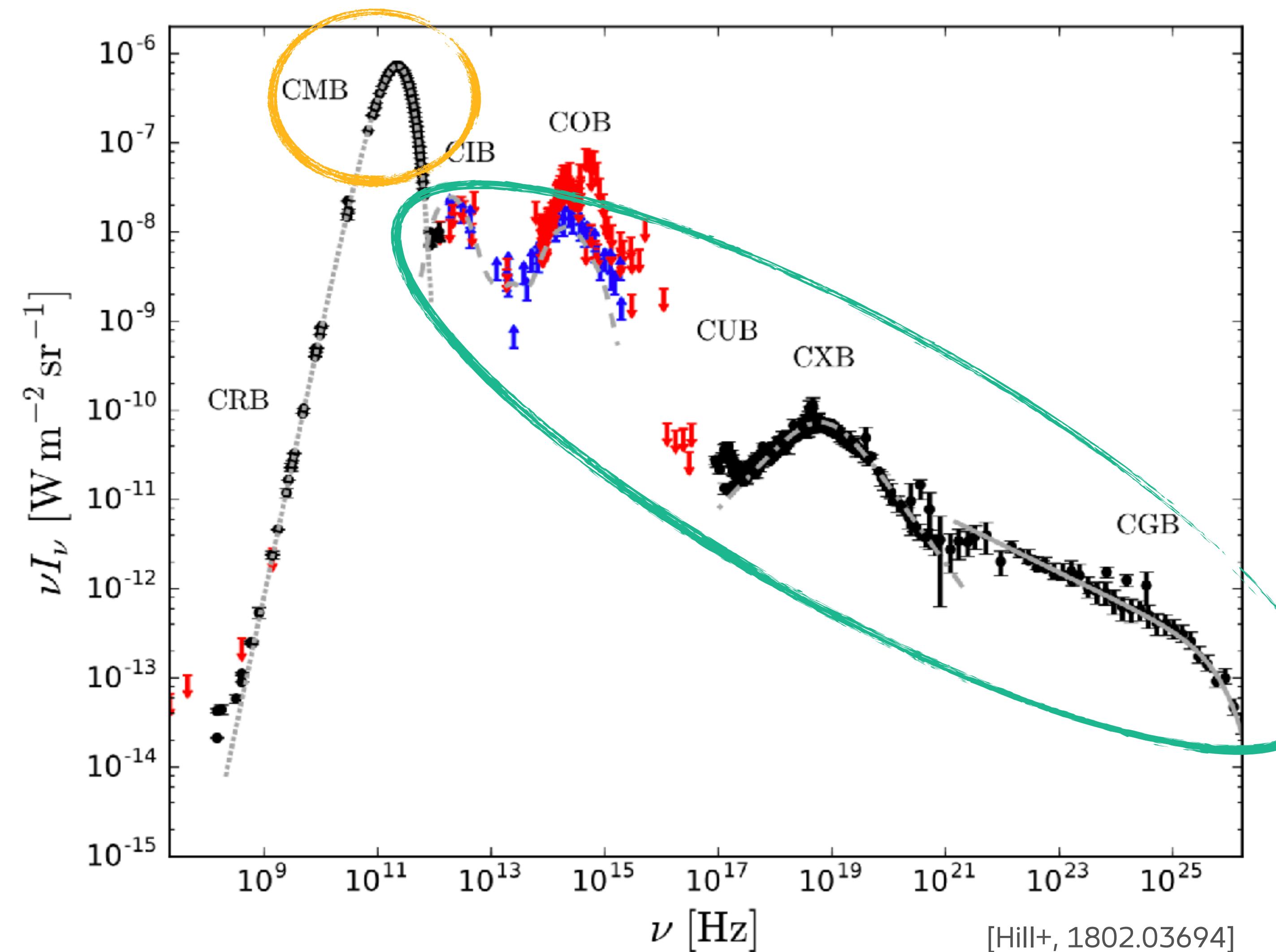
# The observable universe



**The CMB...  
and the CGWB?**

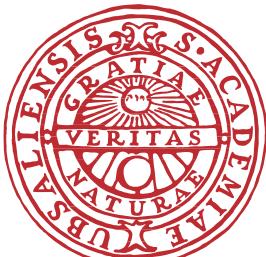
PABLO  
CARLOS  
BUDASSI

# Finding the cosmic microwave background: We were lucky!

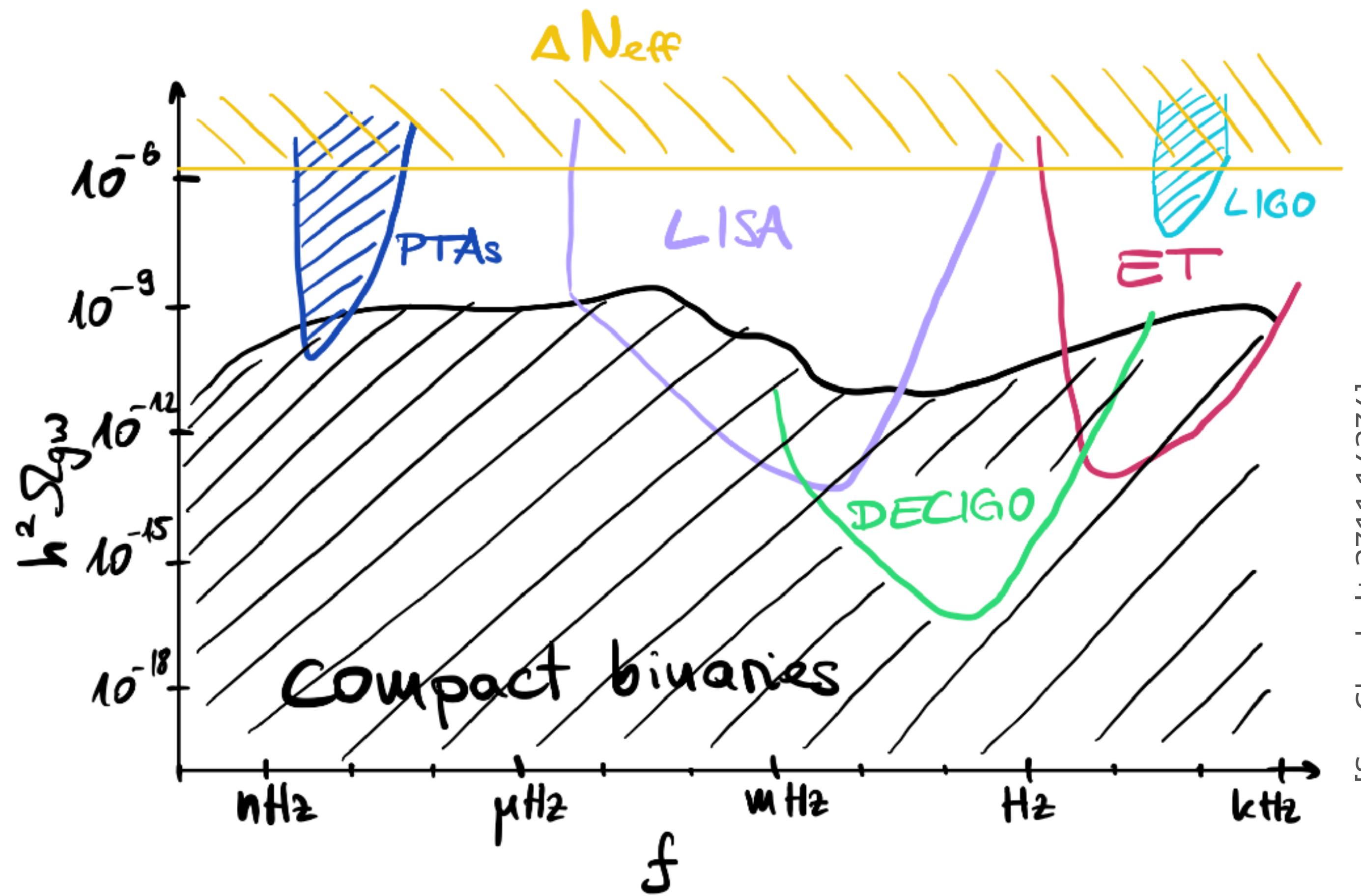


The CMB dominates the photon spectrum!

Astrophysical backgrounds are luckily only active at higher frequencies (or can be masked out)



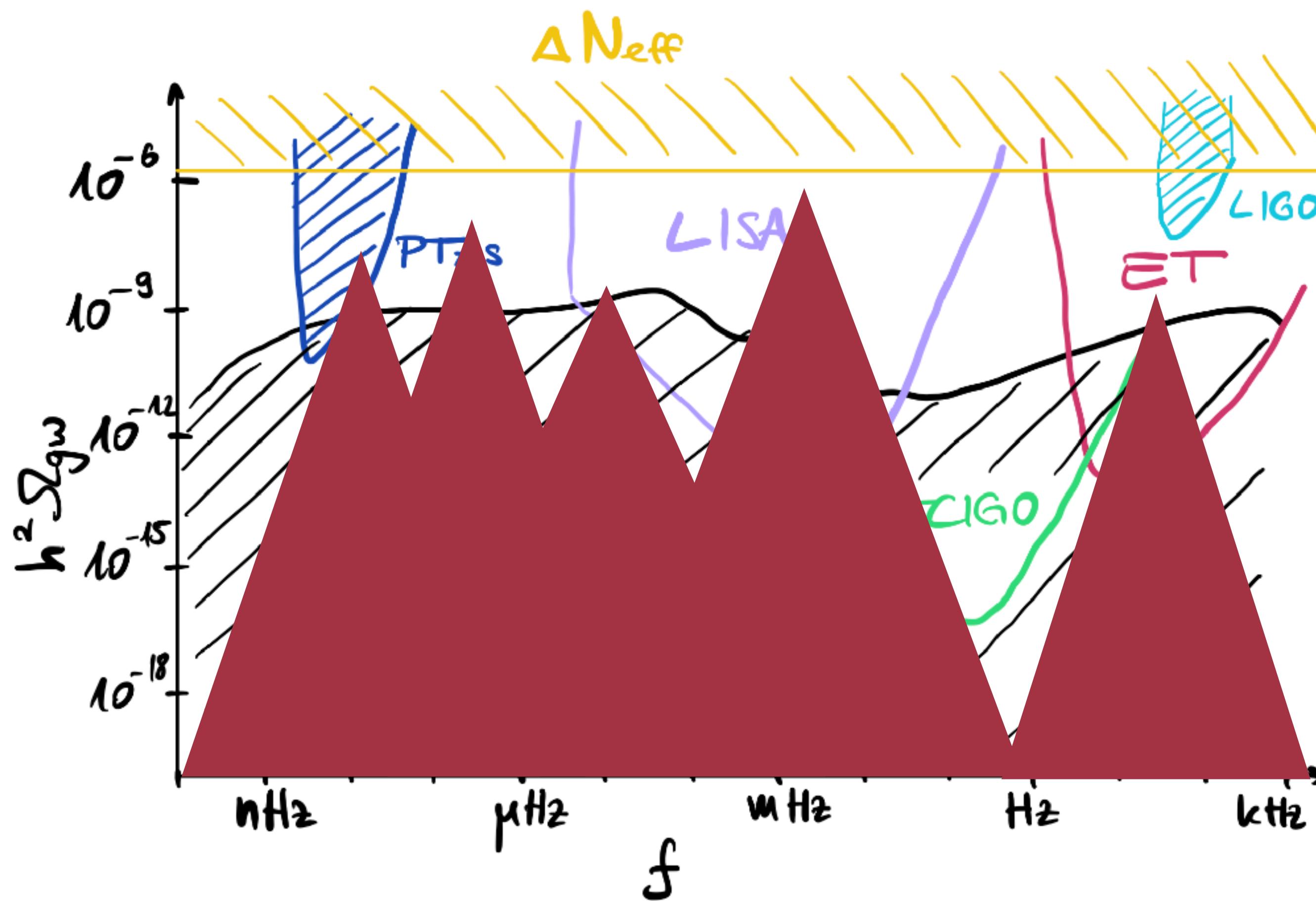
# An attemptive map of the gravitational wave background



[See Ghoshal+ 2311.16236]



# An attemptive map of the gravitational wave background

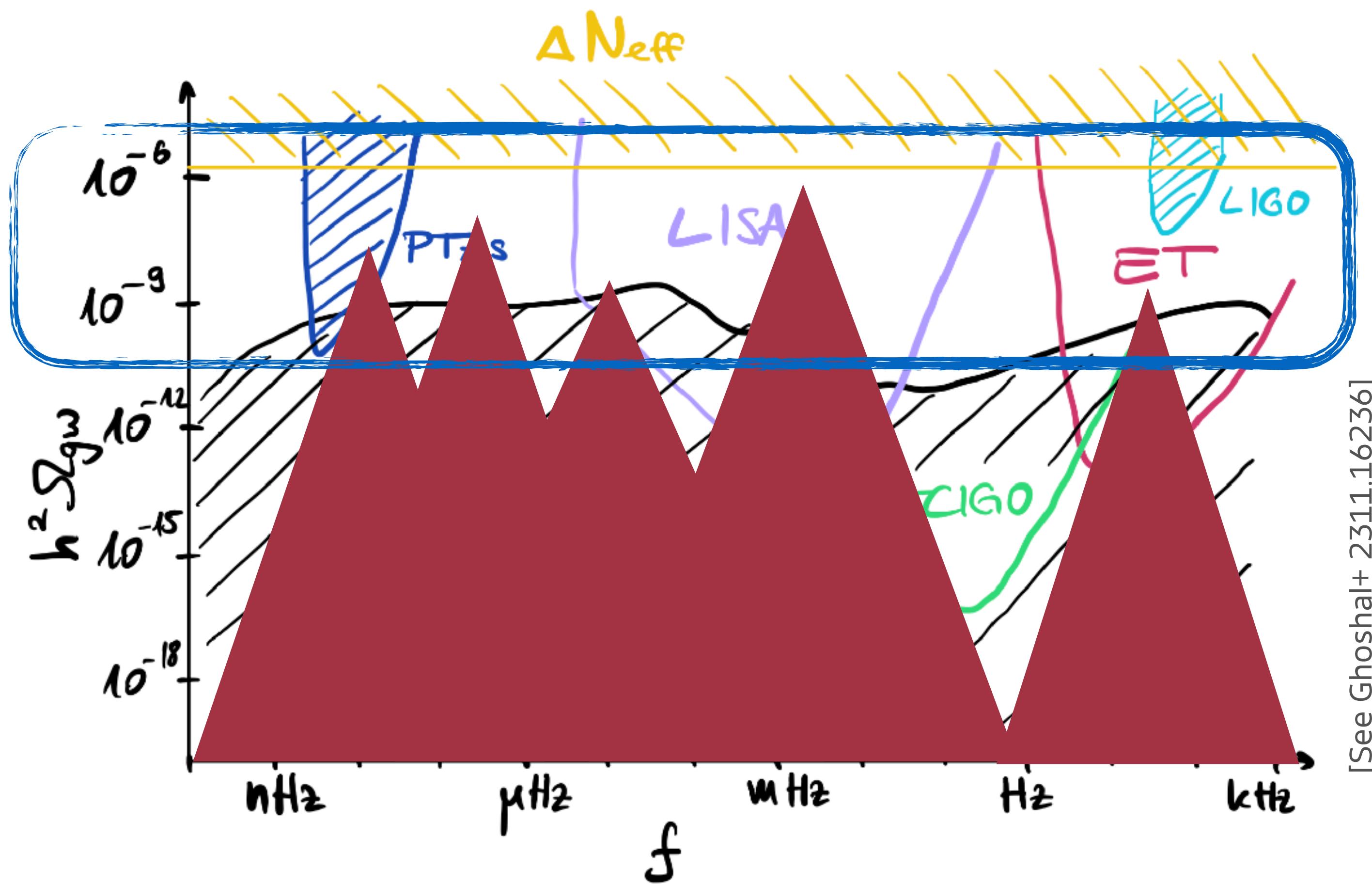


The CGWB would appear as a bump(s) on top of the astrophysical background

[See Ghoshal+ 2311.16236]

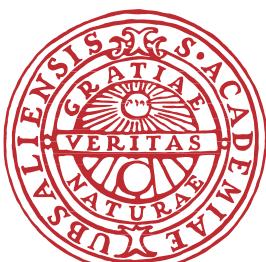


# An attemptive map of the gravitational wave background

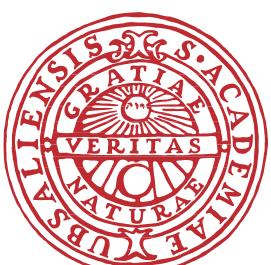
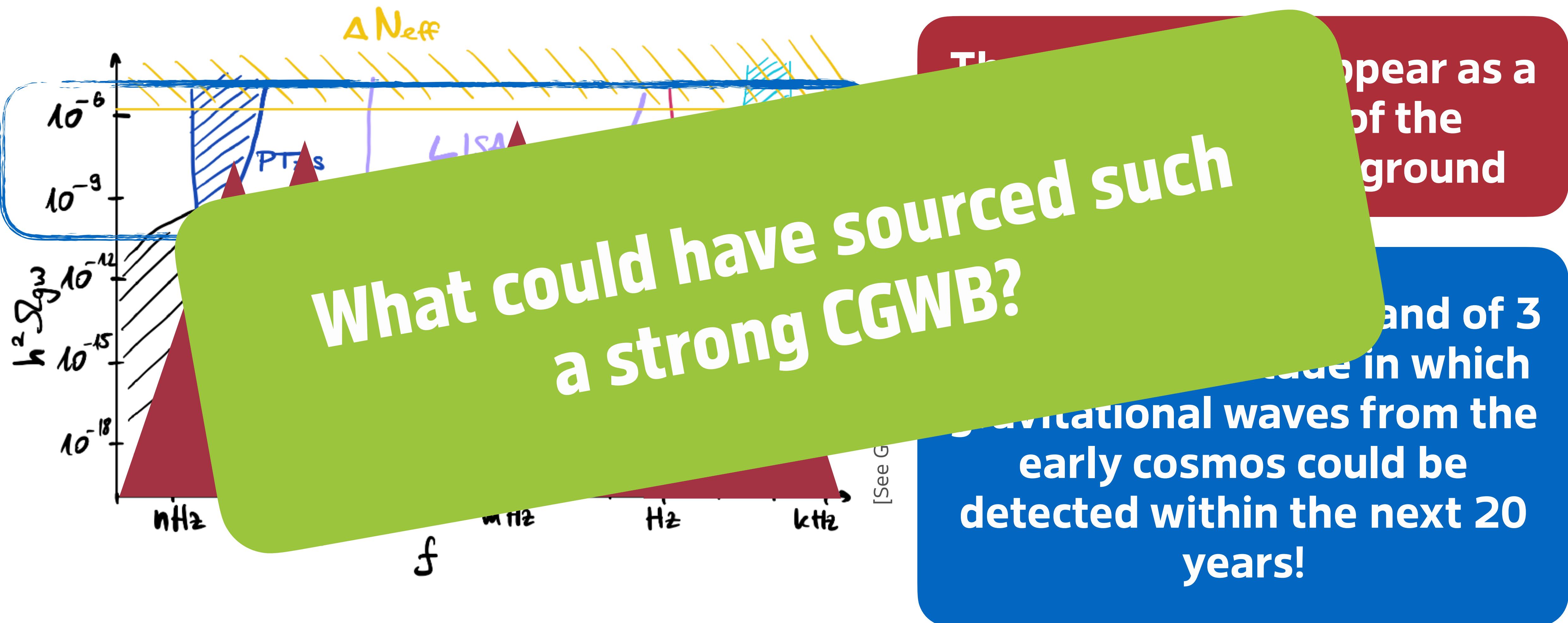


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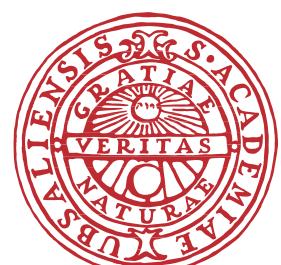
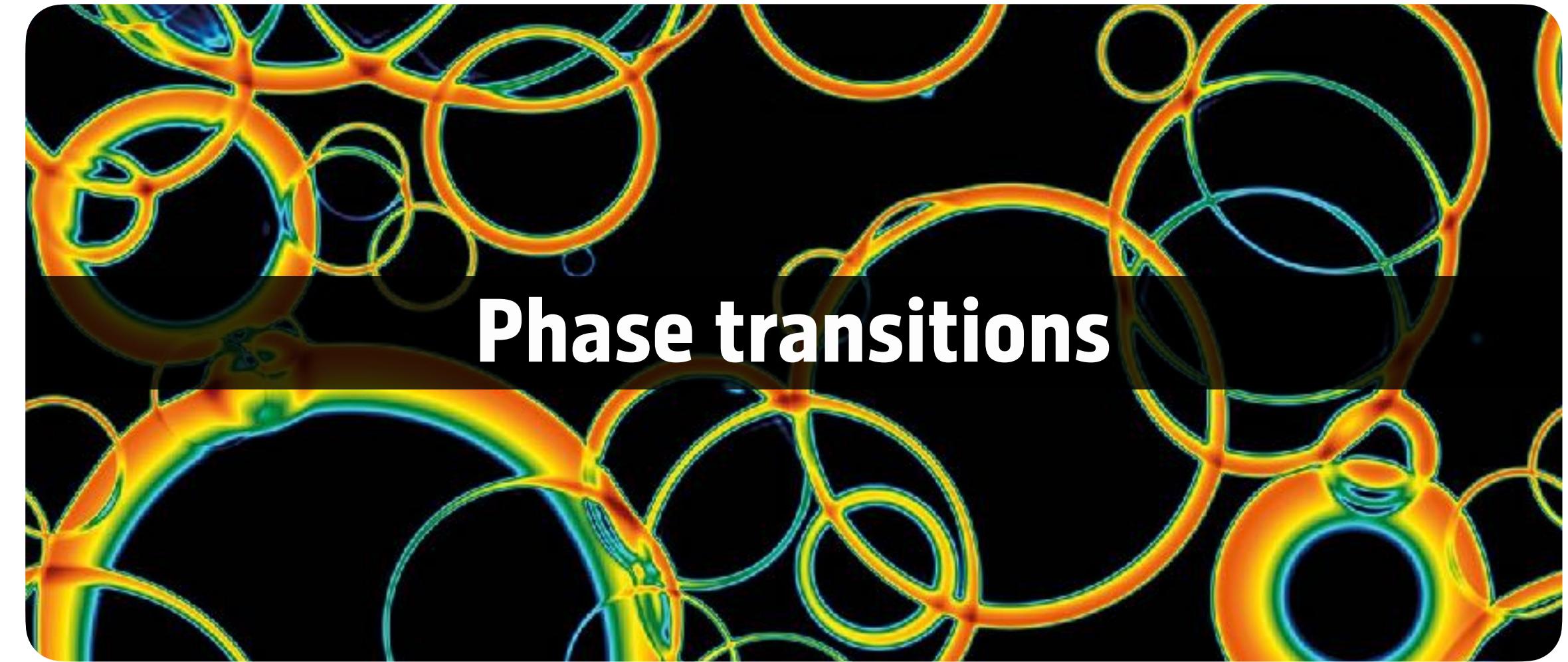
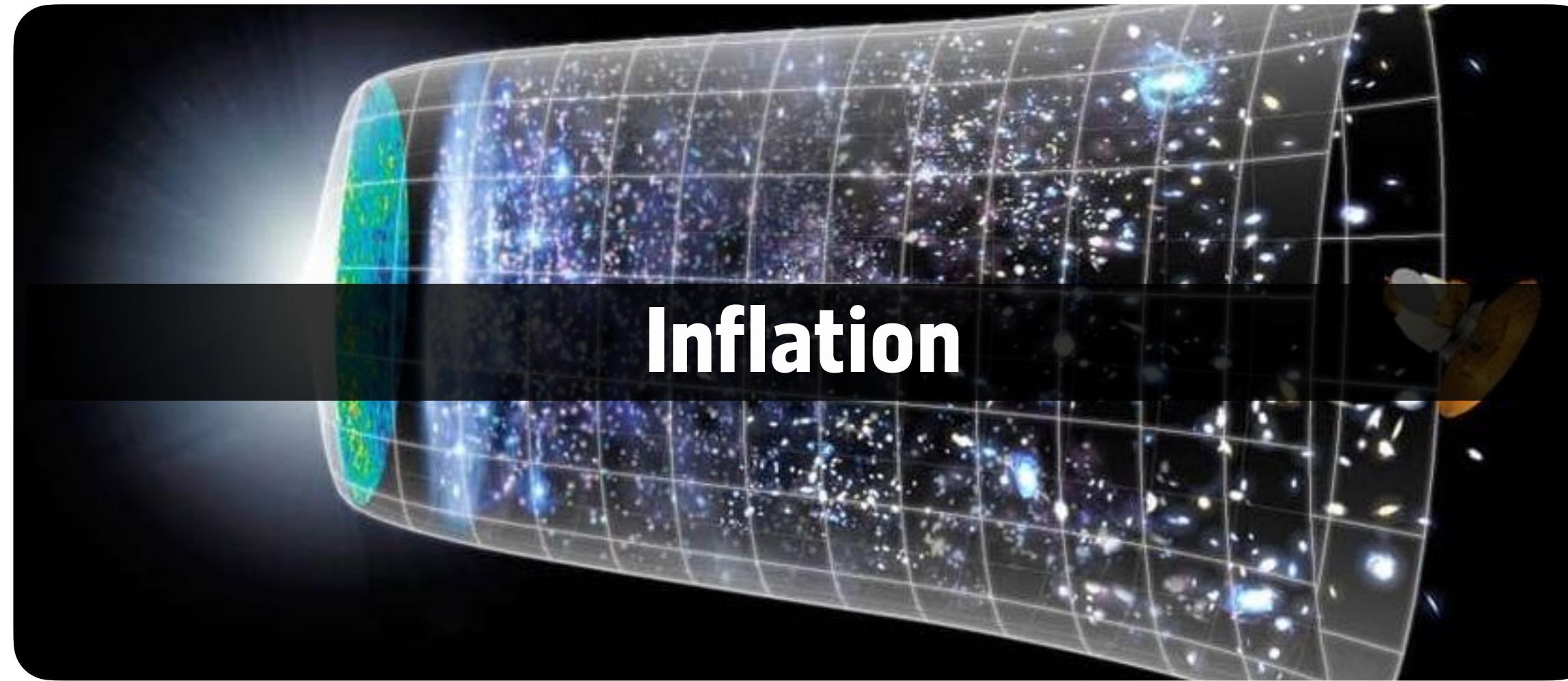
There's just a small band of 3 order of magnitude in which gravitational waves from the early cosmos could be detected within the next 20 years!



# An attemptive map of the gravitational wave background

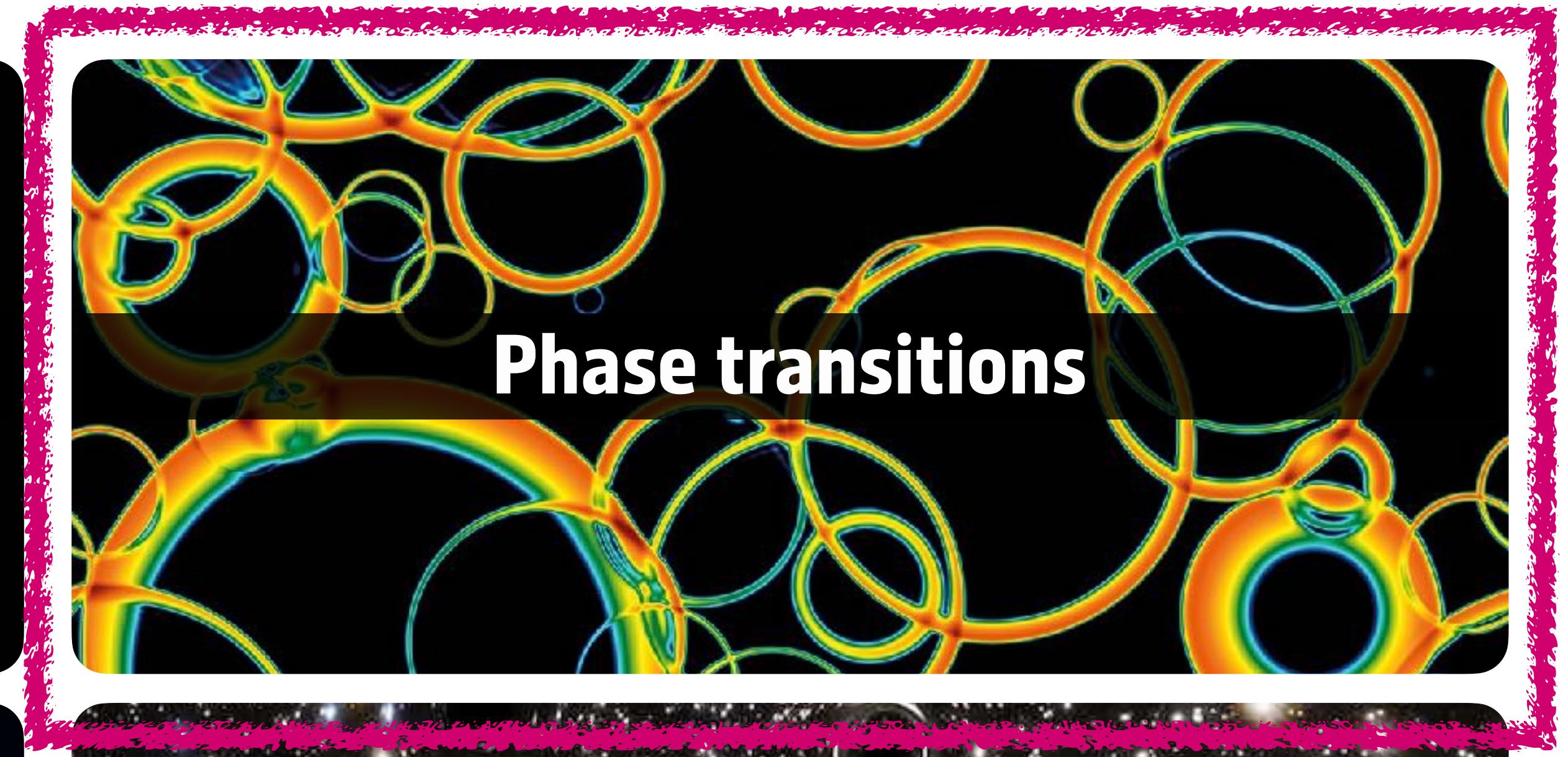
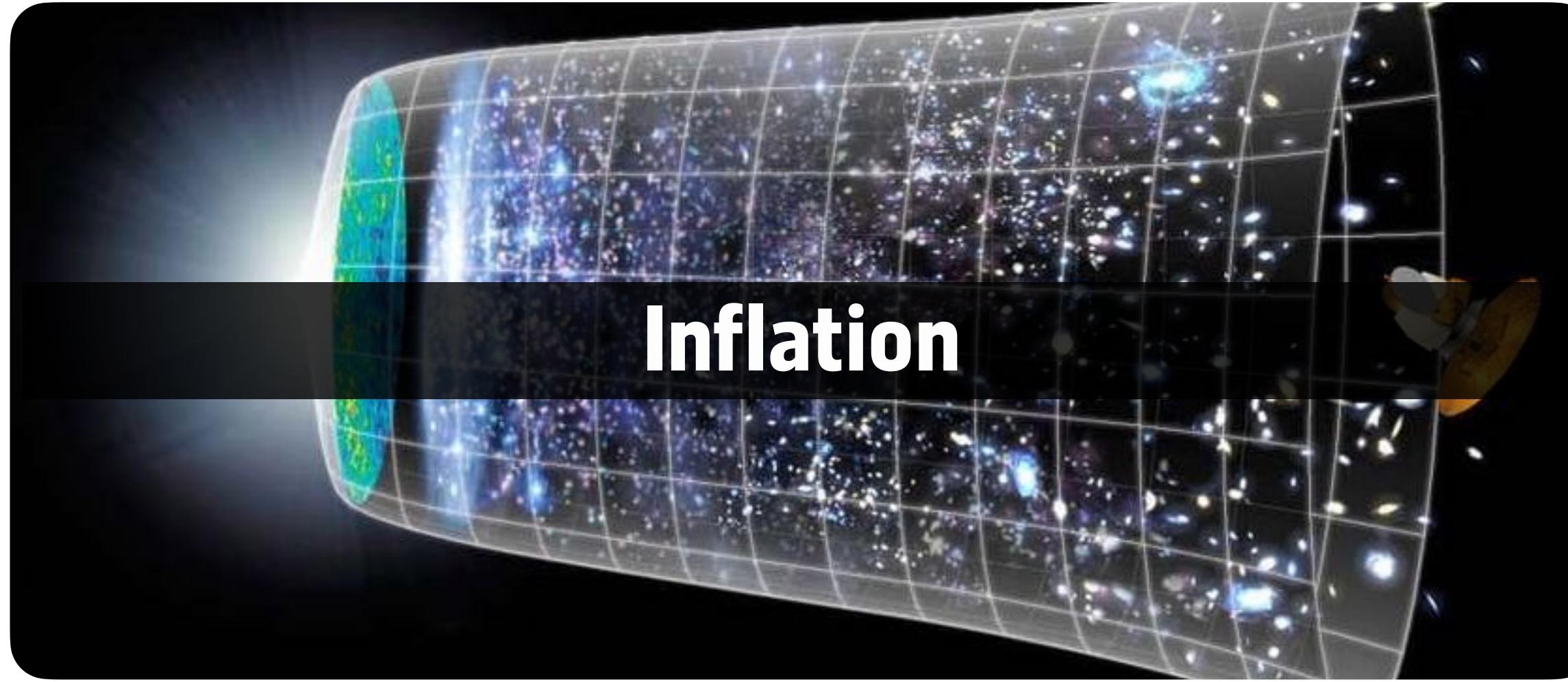


# Possible sources of a loud CGWB

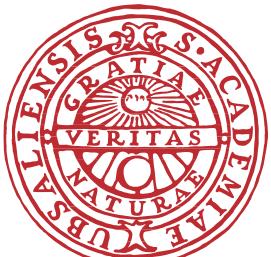
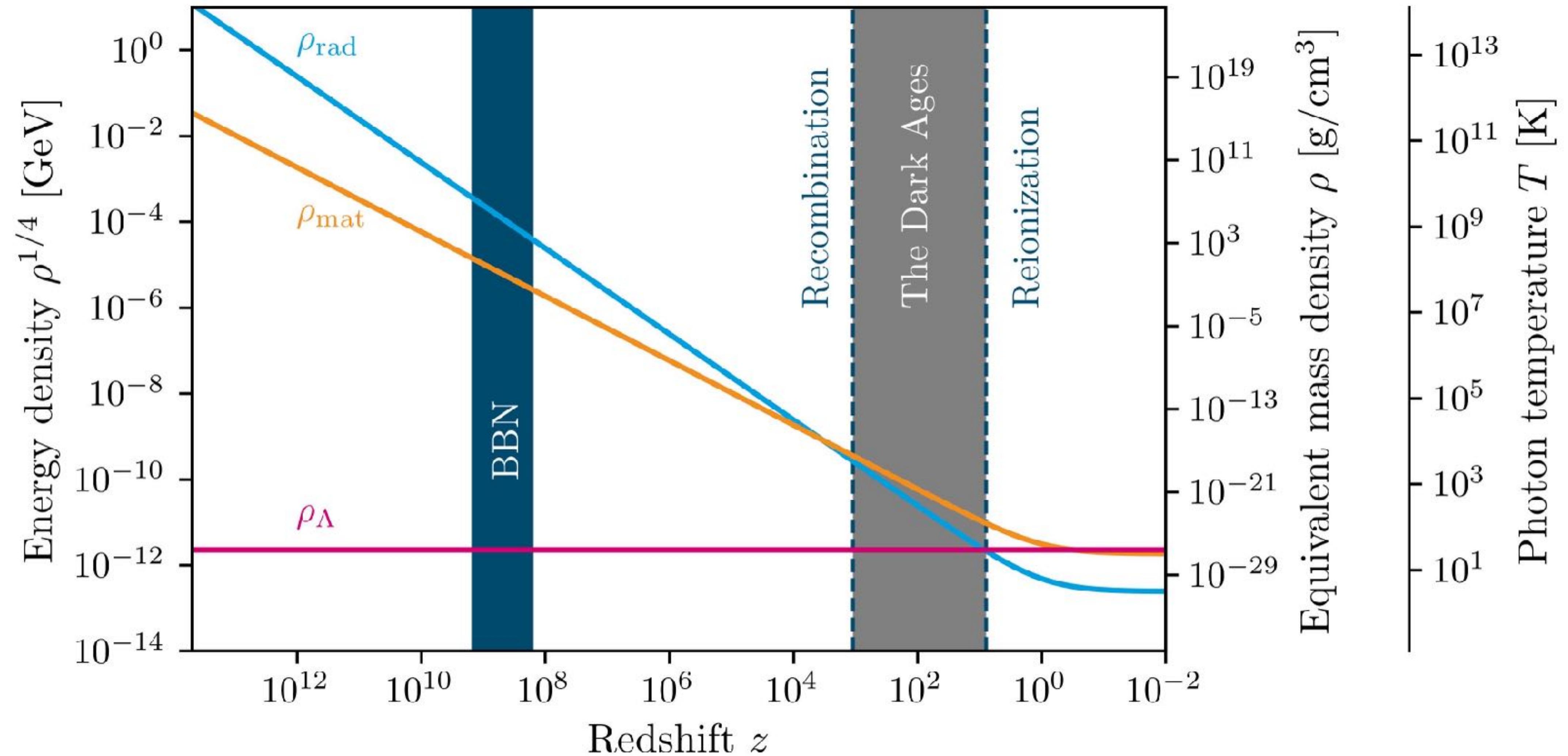


# Possible sources of a loud CGWB

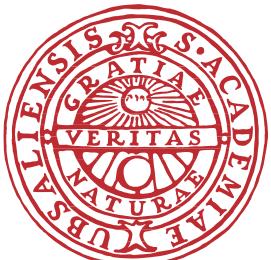
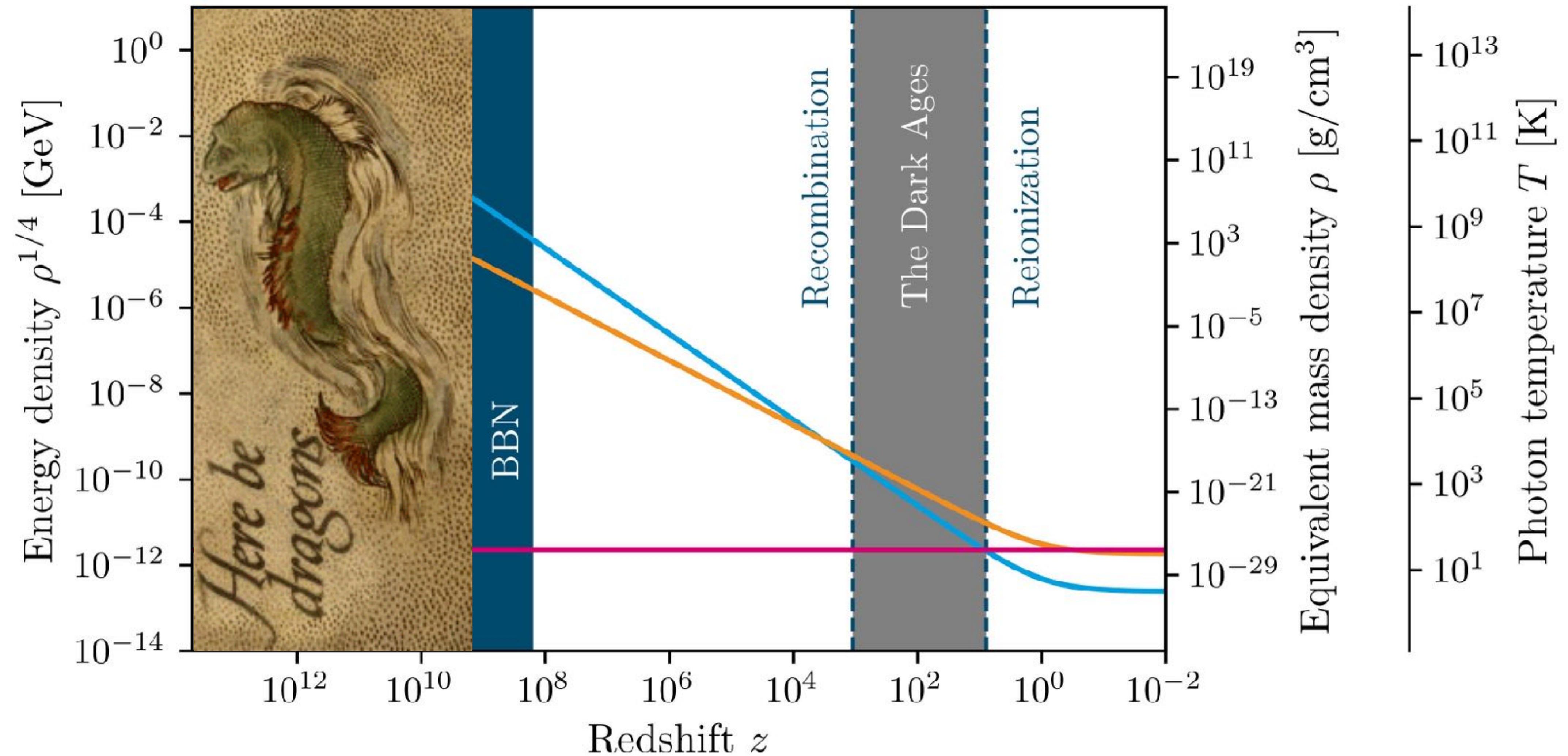
**Today's focus**



# A brief history of time.

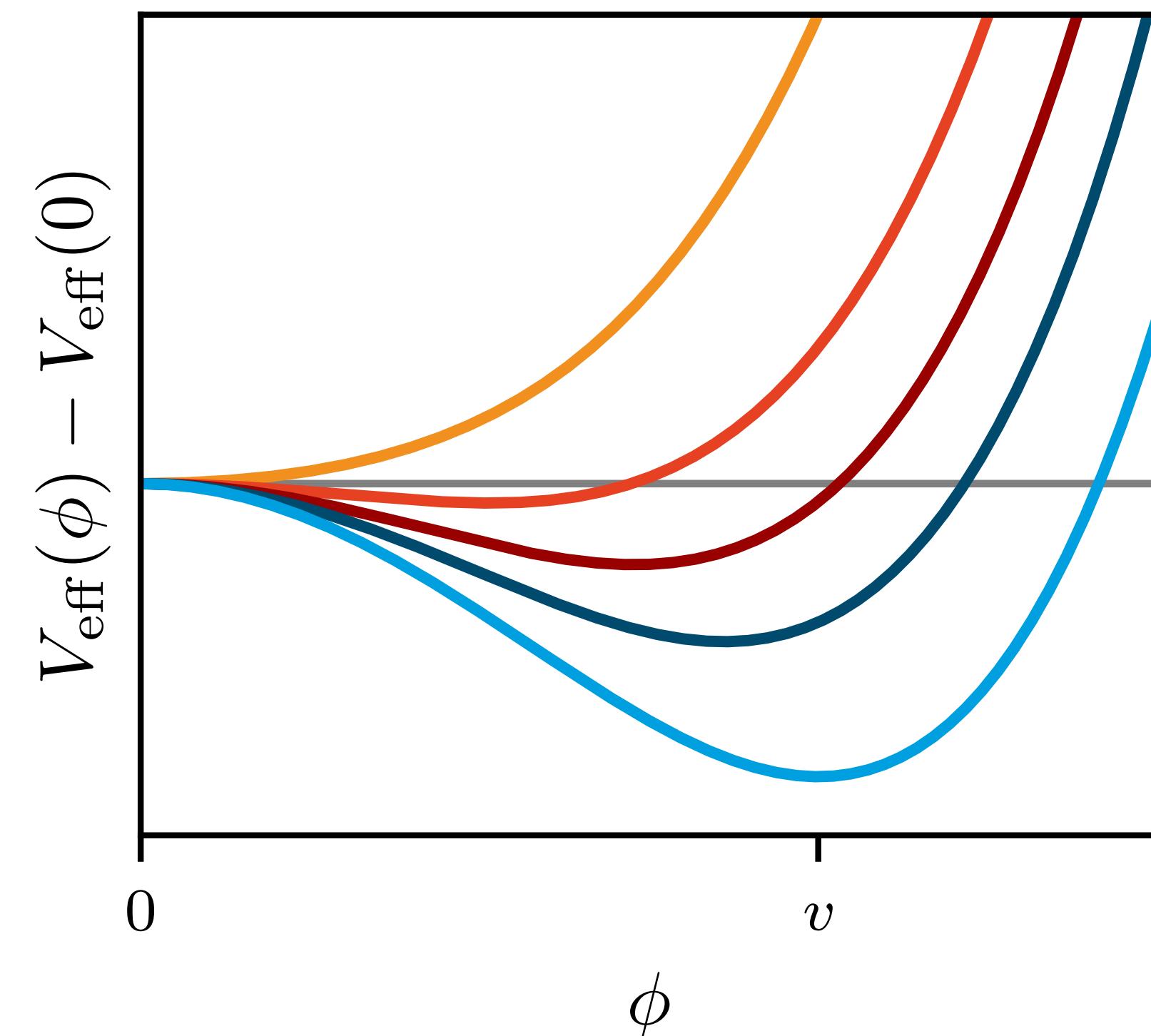


# A brief history of time.



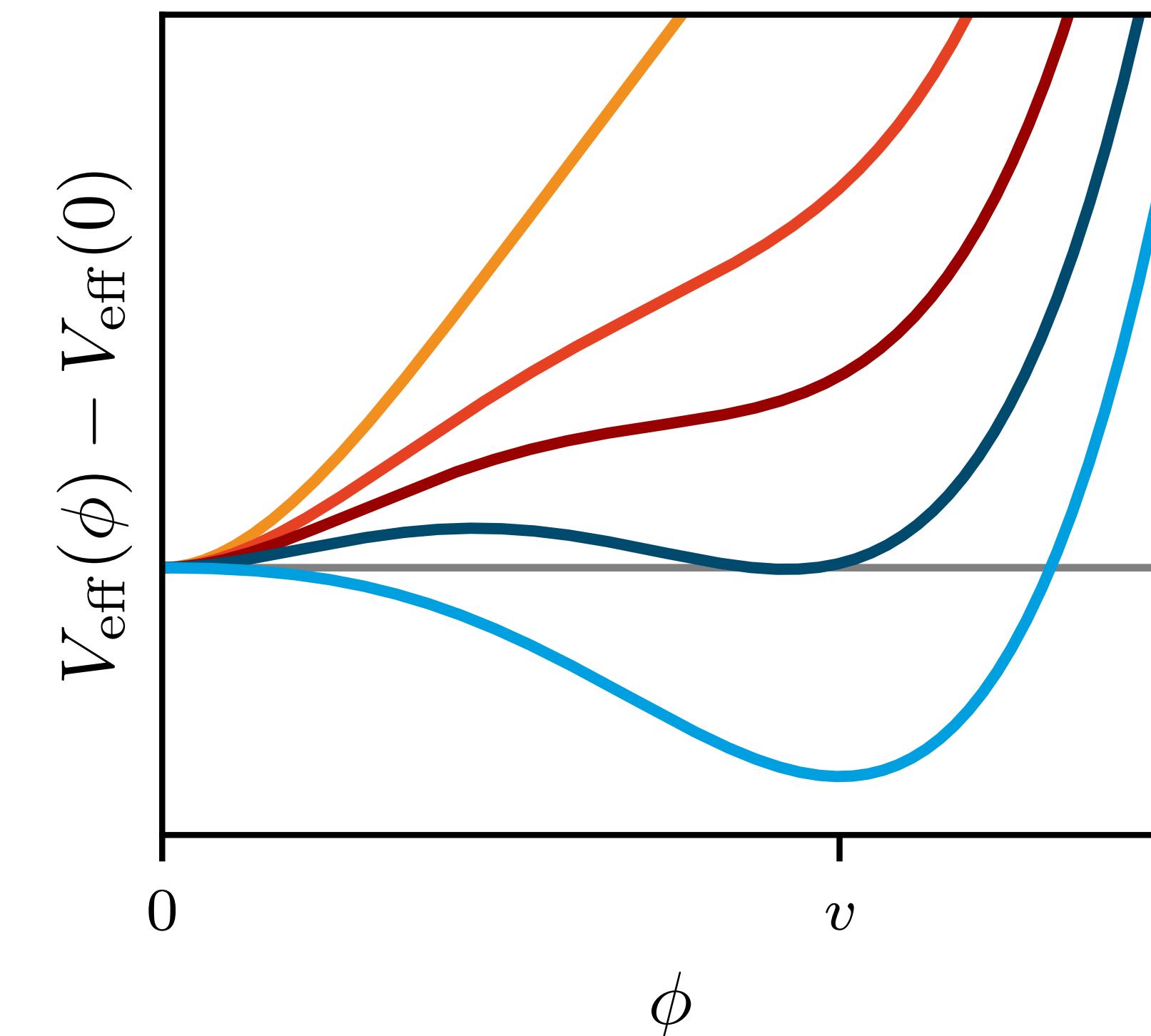
# First-order phase transitions vs. cross-overs

Cross-over phase transition

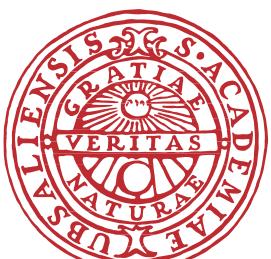


A scalar field “rolls down” from  $\phi = 0$  to  $\phi = v$ , when the plasma cools from **high temperatures** to **low temperatures**.

First-order phase transition

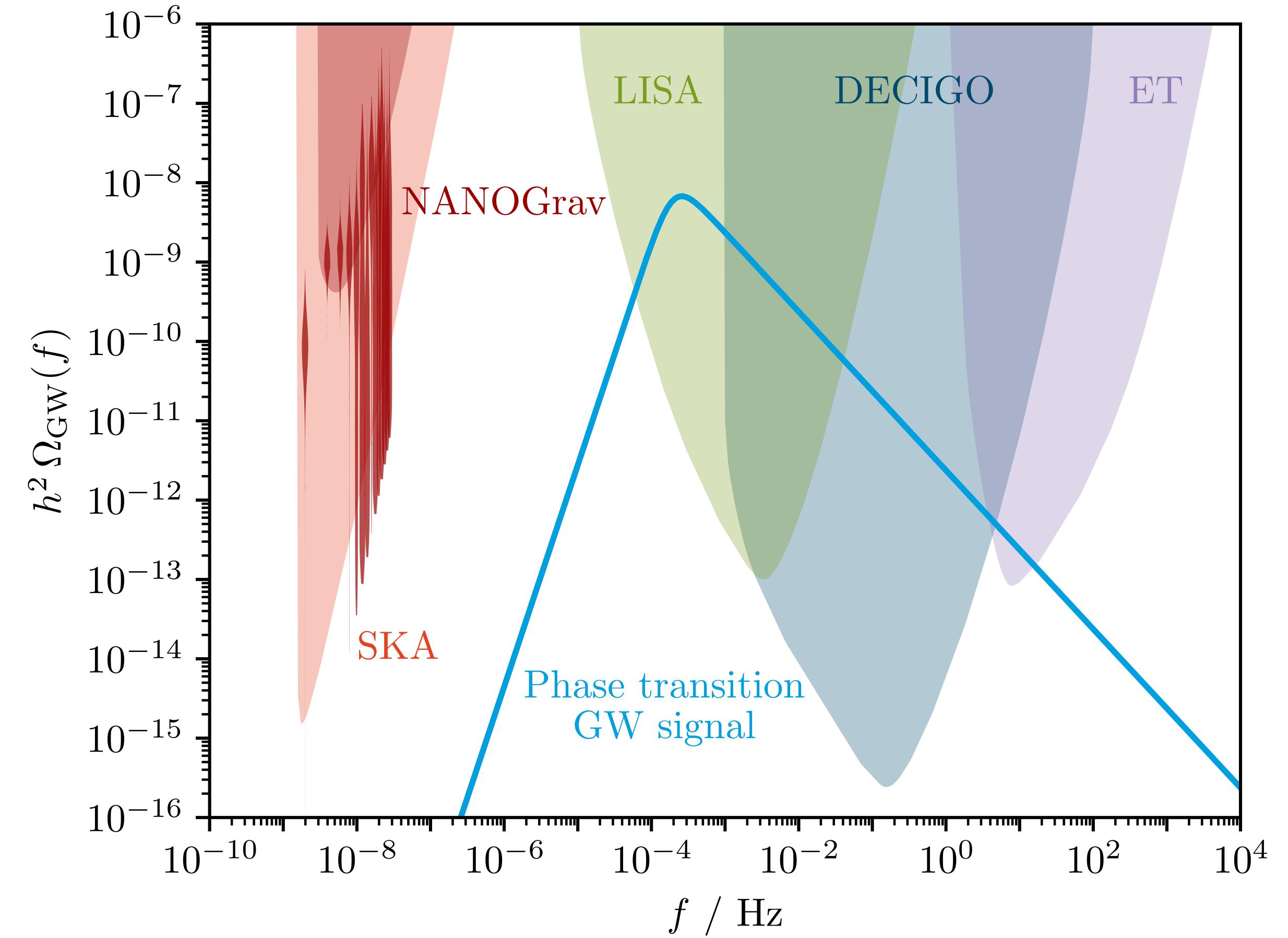
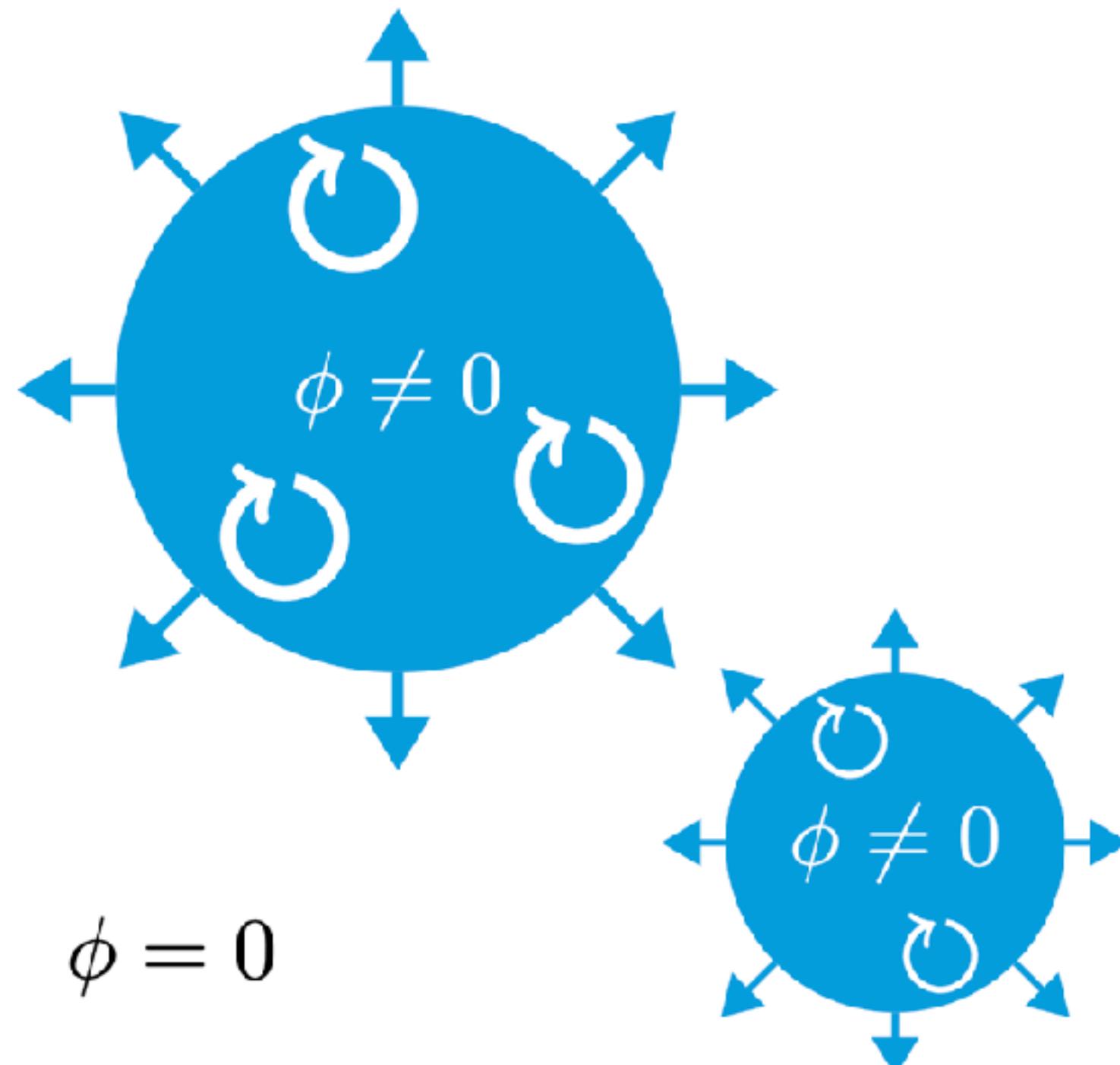


A scalar field tunnels to the true potential minimum  $\phi \neq 0$  to minimize its free energy / its action.

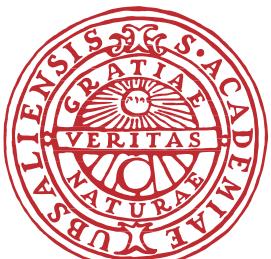


# First-order phase transitions produce GWs

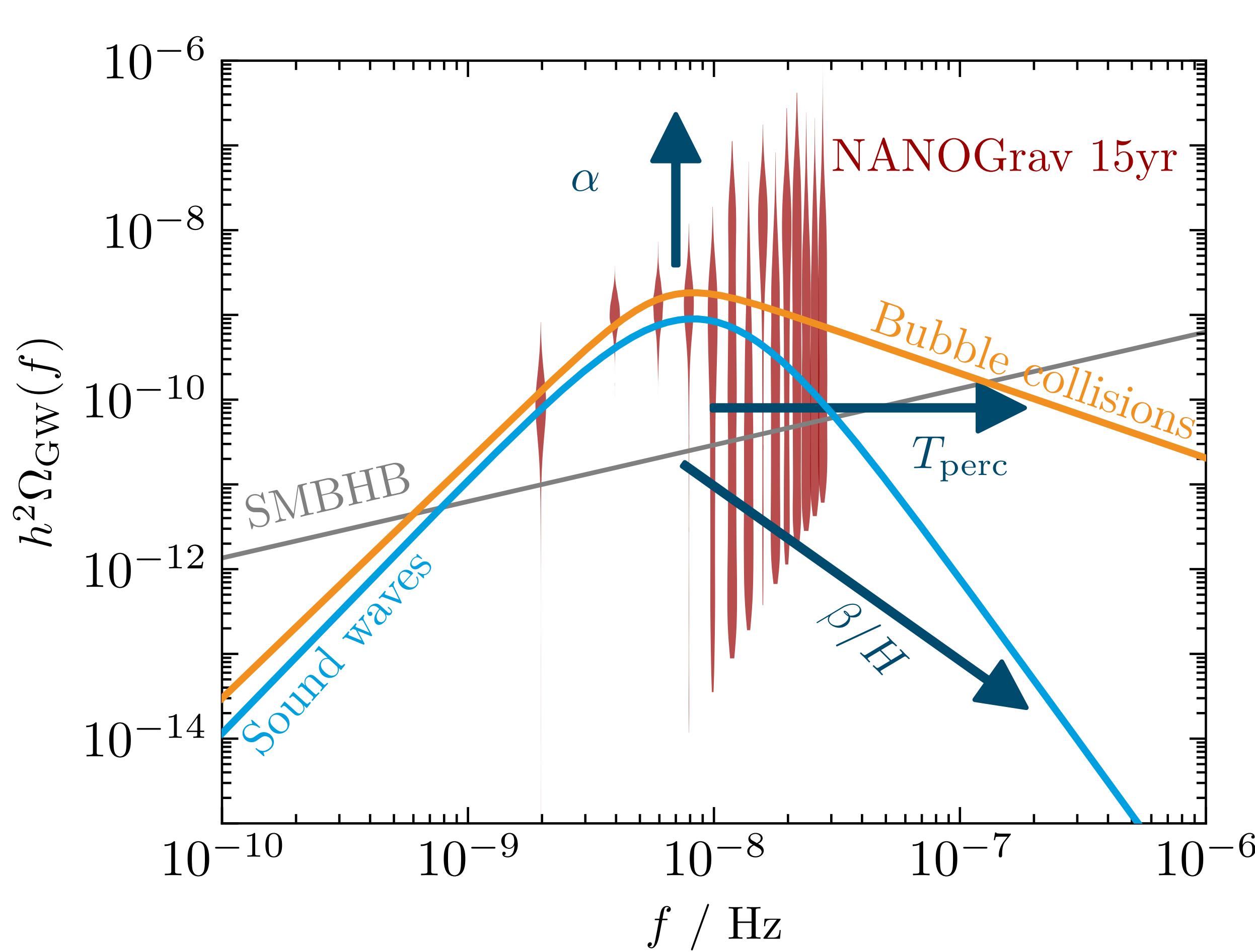
Bubbles of the new phase nucleate, collide and perturb the plasma...



... giving rise to an observable stochastic gravitational wave background.



# How loud can a first-order phase transition be?

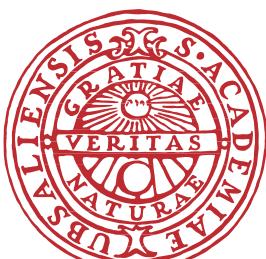


$$h^2 \Omega_{\text{GW}}^{\text{sw}, \text{bw}}(f) \simeq 10^{-6} \left( \frac{\alpha}{\alpha + 1} \right)^2 \left( \frac{H}{\beta} \right)^{1,2} \mathcal{S} \left( \frac{f}{f_{\text{peak}}} \right)$$

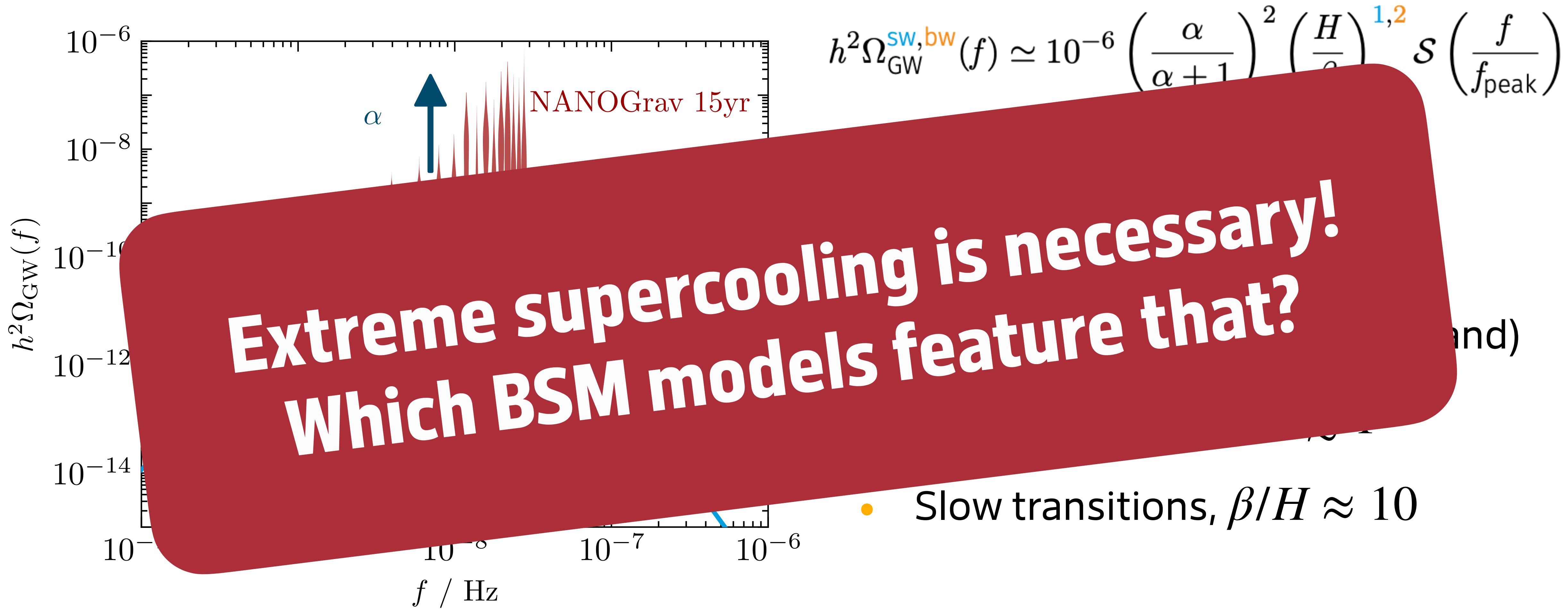
with  $f_{\text{peak}} \simeq 0.1 \text{ nHz} \times \frac{\beta}{H} \times \frac{T}{\text{MeV}}$

In order to be louder than astrophysics (in the PTA & LISA band)

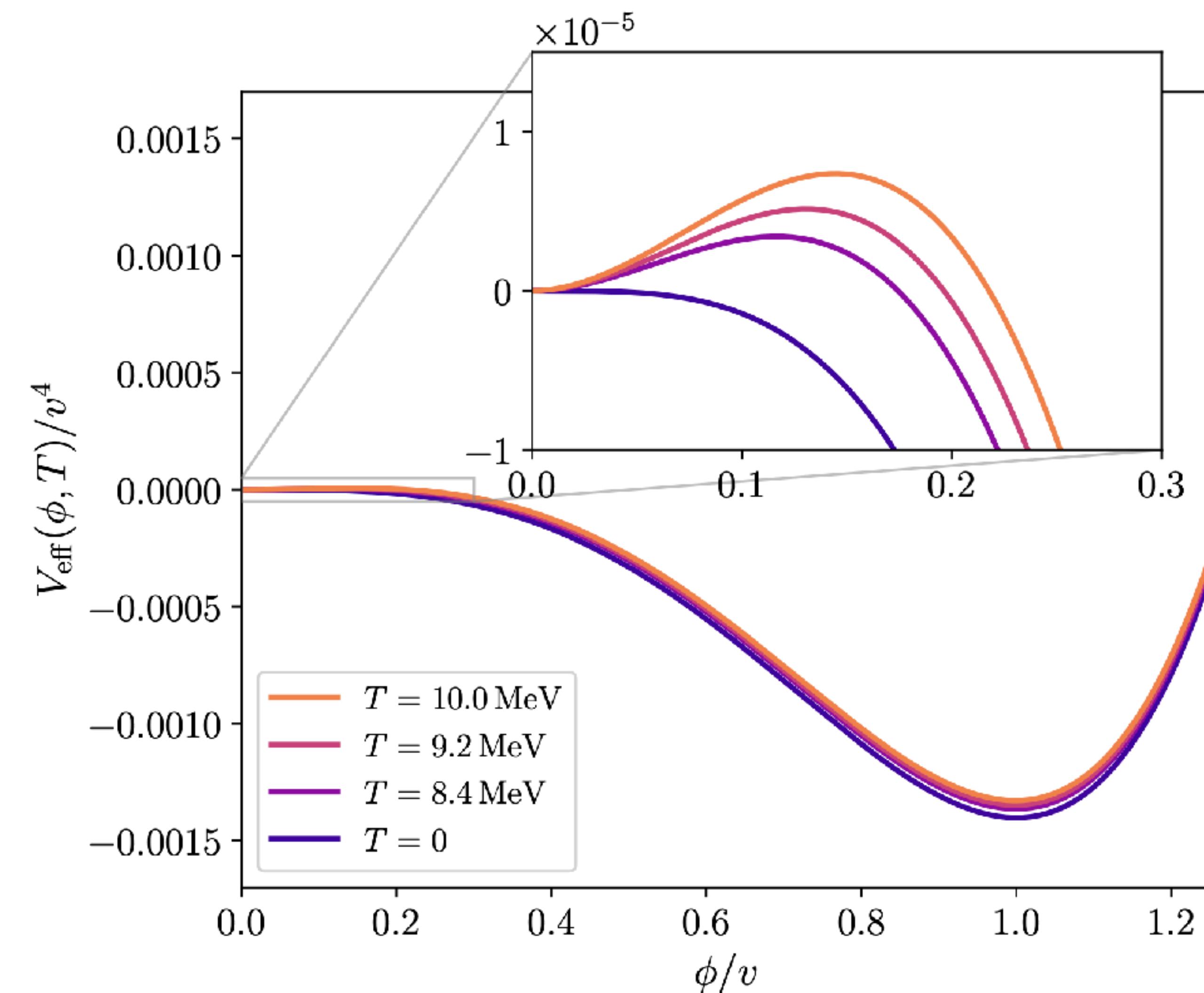
- Strong transitions,  $\alpha \gtrsim 1$
- Slow transitions,  $\beta/H \approx 10$



# How loud can a first-order phase transition be?



# How to generate a potential barrier and make it last for long?



- 1) A generic model allowing for a first-order phase transition **with a lot of fine tuning**

[Ongoing work with Safa and Rikard on dim-6 operators in SMEFT]

- 2) A model with a so-called „vev flip-flop“ and **some fine tuning**

[Ongoing work with André, Harri and Stefano on the (N)2HDM]

- 3) A classically conformal model, **no fine-tuning** needed

[Ongoing work with Torsten Bringmann, Kai Schmidt-Hoberg, Thomas Konstandin, Jonas Matuszak]



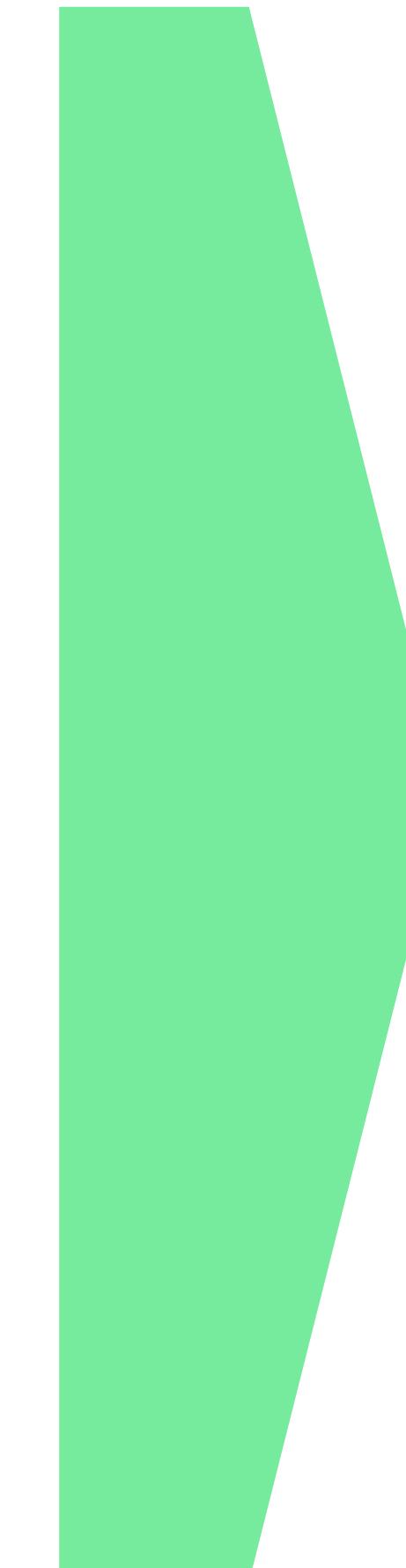
# Reliable computation of the GW spectra for strong supercooling?

There's currently no public code that can deal with strong supercooling!

The available codes are specific to the Electroweak symmetry breaking

Available codes are specialized on specific sub-tasks like phase tracing, the bounce action, the bubble nucleation rate, ...

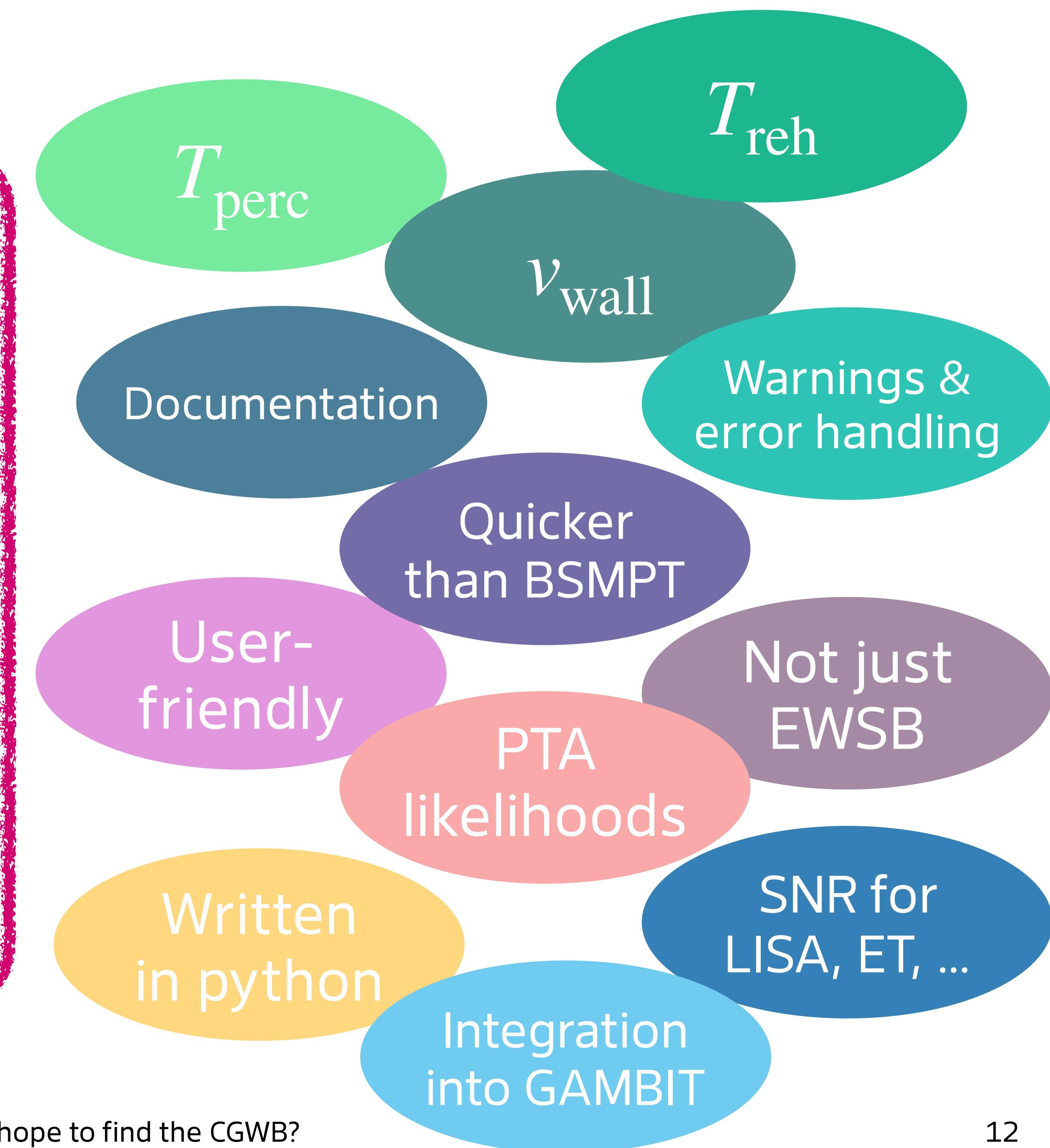
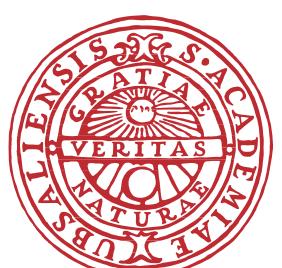
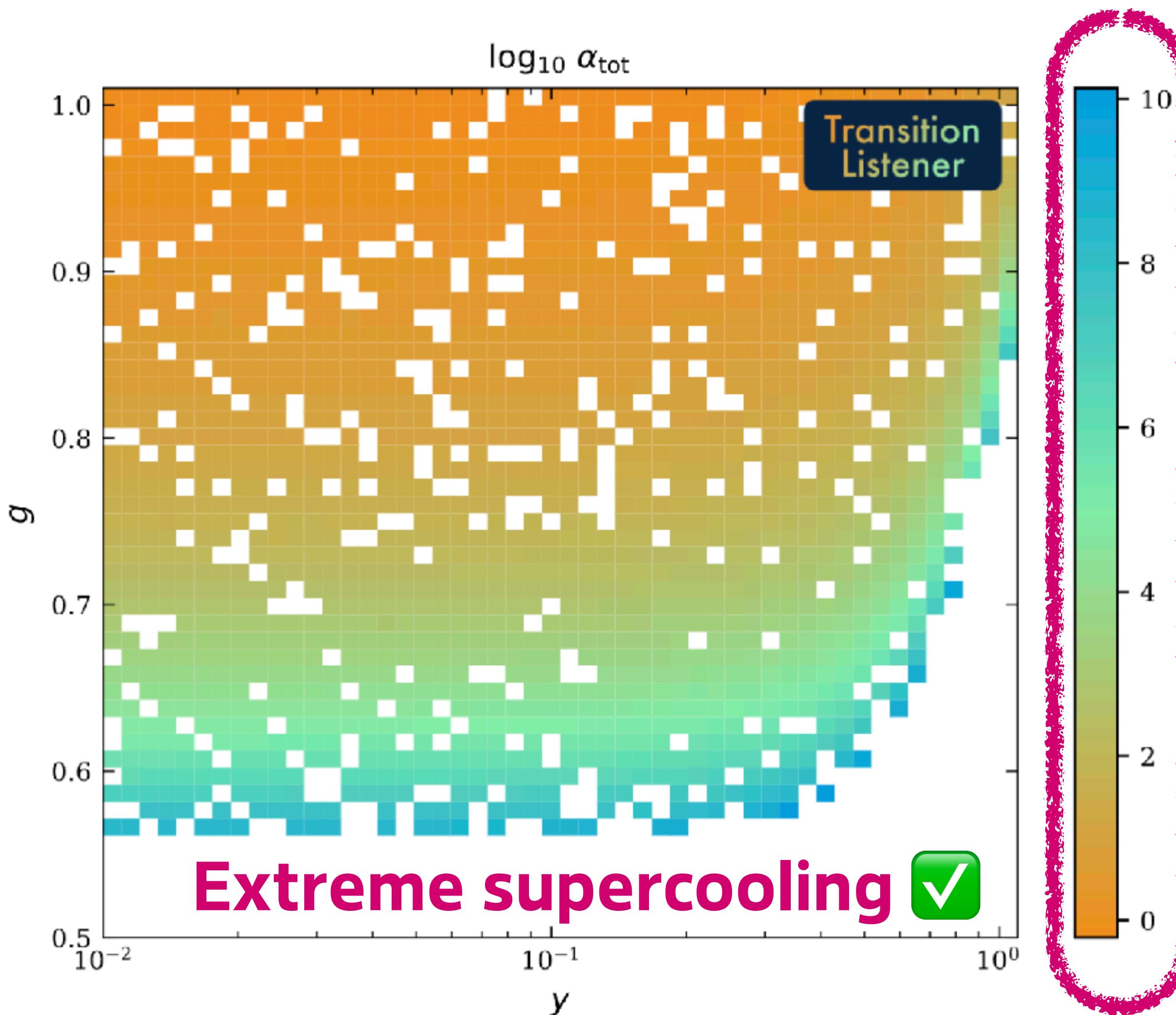
They are not integrated into the ecosystem of global fits, i.e. GAMBIT



[Ongoing work Jonas Matuszak]



# We're on it!



# At Last, There's -

A globe-span-

Astronomers detect 'cosmic bass note' of gravitational waves

Sound comes from the merging of supermassive black holes across the universe, according to scientists

# Gravitational Waves

Scientists have finally 'heard' the chorus of gravitational waves that ripple through the universe

Black Holes in Space

Gravitational wave at the center of the Mi-

# Scientists Re-

come from c-

holes

# The Cosmos Is Thrumming With Gravitational Waves, Astronomers Find

Radio telescopes around the world picked up a telltale hum reverberating across the cosmos, most likely from supermassive black holes merging in the early universe.

It may be

massive black

# of Low-Frequency Gravitational Waves

the waves, w-

ed from pa

# Scientists 'hear' cosmic hum from gravitational waves

ing everything in the universe.

# First Evidence of Giant Gravitational Waves Thrills Astronomers

ture tuning in to a never-before-seen type of gravitational waves that was announced earlier this year

## Monster gravitational

spotted for first

time

in

the

first

time

in

the

first

# For first time ever, scientists "hear" gravitational waves rippling through the universe

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If we care about phase transitions in BSM physics, it is time to start building our CGWB analysis pipelines!

# Summary

- We can indeed hope to find (or even *have found*) GWs from cosmology!
- But: It still remains to be seen if the early universe has produced **stronger GWs than astrophysics**
- We should put our emphasis in the **study of the strongest GW sources!**
- There are **three model classes** which can give strong GWs from phase transitions with varying levels of fine-tuning
- Soon: **TransitionListener v2** for studying all these models and comparing with data



# Backup slides